16.2 Structure of the atom (full E.1, SL)

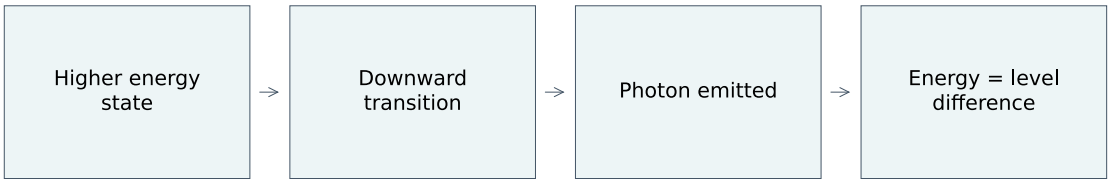
Name: _____ Class: _____ Date: _____ **Total: 22 marks**

KEY WORDS nucleon number • proton number • emission spectrum

Objective

Practise the full guide E.1 toolkit at standard level: the Geiger–Marsden–Rutherford experiment and the discovery of the nucleus, nuclear notation ${}^{A}_{Z}X$, emission and absorption spectra as evidence for discrete energy levels, photon emission and absorption in transitions with $E = hf$, and spectra as fingerprints of chemical composition. Original practice; not an official IB paper.

1 Key knowledge and worked examples



■ **The E.1 rule set (SL)**

$$E = hf \qquad {}^{A}_{Z}X$$

Alpha scattering: most particles pass through (empty atom), few deflect far (tiny dense positive nucleus). Spectral lines correspond one-to-one with transitions between discrete levels; the set of lines identifies the element.

■ **Worked example: reading a line**

Levels at -1.5 eV and -3.4 eV: a downward jump emits $E = 1.9\ \text{eV} = 3.0\times10^{-19}$ J, so $f = E/h = (3.0\times10^{-19})/(6.6\times10^{-34}) = 4.6\times10^{14}$ Hz. The atom absorbs exactly this frequency climbing up – emission and absorption lines match. Two habits: convert eV to joules before using $E = hf$, and remember each line is one gap, not one level.

2 Practice

2.1 State what the Geiger–Marsden–Rutherford experiment revealed about the structure of the atom, citing two observations. [3]

2.2 Write the meaning of each symbol in nuclear notation ${}^{A}_{Z}X$. [2]

2.3 Why do emission and absorption spectra consist of sharp lines rather than continuous bands? [2]

2.4 State one use of spectral analysis that identifies chemical composition at astronomical distances. [1]

3 Exam-style questions

3.1 An atom has energy levels at -1.5 eV, -3.4 eV and -13.6 eV. Calculate the frequency of the photon emitted in the $-3.4 \rightarrow -1.5$ eV transition, and the energy (in eV) of the photon that ionizes the atom from the -1.5 eV level. [4]

3.2 Explain why almost all alpha particles passed straight through the gold foil while about 1 in 8000 bounced back, and what each observation implies. [4]

3.3 A gas of element X shows absorption lines at exactly the same frequencies as its emission lines. Explain how this identifies X in the atmosphere of a distant star. [3]

3.4 The nuclide $^{235}_{92}U$ is used in reactors. State its proton number, nucleon number, and the number of neutrons in its nucleus. [3]